

Erratum: MK Classifications for F- and G-Type Stars. II*†

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Santa Cruz, California*MK spectral classifications are given for 166 stars of HD types F2–G5, and having $m_V=7.5$ or brighter. The classifications were made on slit spectrograms of dispersion 75 \AA/mm .

IN Paper I of this series, Harlan (1969) presented MK classifications for 314 stars having HD spectral types between F2 and G5, and $m_V=7.5$ or brighter. The classifications were made on prismatic spectrograms of dispersion 75 \AA/mm at $H\gamma$, obtained at the 36-inch refractor. This second list of types for 166

TABLE I. MK classifications of F and G stars.

HD	α (1900) δ	m_V	HD	Type MK	HD	α (1900) δ	m_V	HD	Type MK
417	0 ^h 03 ^m 7 +24°54'	6.4	G5	G8 II-III	97561	11 ^h 08 ^m 5 +20°41'	6.9	G0	G5 III
483	04.2 +16 59	7.2	G0	G2 III	98354	13.7 +14 49	6.6	G0	F9 V
1375	12.8 +12 13	6.6	G5	G8 II	100180	26.6 +14 55	6.2	G0	G0 V
2302	21.8 +24 29	6.7	F5	F7 IV	101484	35.6 +21 54	5.4	G5	K1 III
2925	27.4 +22 38	7.0	G0	G8 III-IV	101688	37.0 +22 46	6.6	F2	F3 V
2954	27.7 +15 38	6.9	F5	F6 III	104452	56.6 +22 39	6.6	F8	G0 II
4432	41.5 +21 05	7.0	F8	F8 IV	104883	59.6 +16 51	7.5	F8	F5 III
5036	47.2 +20 52	7.2	F5	F7 V	106022	12 06.9 +29 06	6.4	F2	F5 IV
5137	48.2 +28 58	6.7	G5	G5 III	106057	07.1 +21 06	5.7	G5	G8 III
6566	1 01.4 +13 22	7.3	F2	F6 V	107146	14.1 +17 07	7.0	G5	G2 V
8815	1 21.8 +29 16	7.2	F2	F6 V	107213	12 14.5 +28 43	6.3	F5	F8 V
9986	32.4 +11 35	6.9	G0	G5 V	107569	16.7 +22 25	7.4	F5	F8 III
10995	42.9 +16 32	7.3	G0	G2 IV	111066	41.6 +24 41	6.7	F5	F8 V
11592	48.7 +10 08	7.0	F5	F7 V	111395	43.9 +25 24	6.4	G5	G5 V
13382	2 05.7 +20 55	7.2	G0	G5 V	112060	48.6 +20 02	6.6	G5	G5 IV
13825	09.6 +23 50	6.9	G5	G8 IV	112196	49.7 +22 39	7.0	F8	F8 V
16150	30.5 +13 20	7.4	F5	F6 III	113022	55.7 +18 55	6.1	F5	F7 IV
18803	56.4 +26 14	6.7	G0	G8 V	113095	56.2 +17 40	6.0	G5	K0 III
19270	3 00.9 +12 48	5.8	G5	K3 III	113848	13 01.4 +21 42	6.0	F5	F4 V
19504	03.0 +10 25	7.4	F5	F4 V	114637	06.7 +22 28	6.8	G5	G5 III
23257	3 38.6 +27 37	7.0	G0	G2 IV	114793	13 07.7 +19 17	6.6	G5	G8 III
25230	55.3 +19 55	6.8	G5	K1 III	115782	14.3 +29 07	7.0	F8	F8 IV
25296	56.1 +27 52	7.5	G5	G8 II	117176	23.5 +14 19	5.2	G0	G5 V
25627	59.0 +17 15	6.8	G5	K2 III	117876	28.1 +24 52	6.2	G5	G8 III
33185	5 03.4 +29 40	6.6	F8	F8 V	118244	30.5 +23 01	7.0	F5	F5 V
33252	03.8 +27 26	6.9	F5	F8 V	119126	36.3 +23 01	5.8	G5	K0 III
35956	23.2 +12 29	6.8	F8	G0 V	120064	42.1 +26 12	5.9	F5	F7 III-IV
35984	23.4 +29 07	6.2	F5	F6 III	120510	44.8 +13 30	6.7	F5	F8 V
39156	45.5 +11 28	6.9	G5	K0 IV	120865	47.0 +12 27	6.8	F5	F5 III
39881	50.3 +13 56	6.5	G5	G5 IV	121560	51.1 +14 34	6.2	F5	F8 V
40020	5 51.2 +11 30	6.1	G5	K2 III	122563	13 57.6 +10 11	6.1	G0	F8 IV
41304	59.2 +14 24	6.7	F5	F6 V	122742	58.6 +11 16	6.4	G5	G8 V
44926	6 19.4 +23 30	6.8	G5	K1 III	123845	14 04.9 +16 05	6.7	F5	F7 V
48805	40.1 +23 29	6.5	G5	G8 V	124679	09.9 +10 35	5.4	G5	K1 III
50482	48.4 +21 17	6.6	G5	K1 III	125040	11.9 +20 35	6.4	F5	F8 V
55130	7 06.6 +27 26	6.4	F5	F8 V	126246	19.3 +11 42	6.7	F5	G0 V
56329	11.5 +14 02	7.4	F8	F6 IV	127539	26.9 +18 05	7.2	F5	F5 III
78418	9 03.0 +27 03	6.0	G5	G5 IV	127740	28.0 +14 31	7.0	F5	F5 III
78715	04.6 +22 24	6.1	G5	G8 III	129132	35.8 +22 24	6.2	F5	G0 V
80536	15.3 +25 36	7.3	G0	G2 V	129336	36.9 +12 05	5.6	G5	G8 III
80654	9 15.9 +13 33	6.6	F5	F8 V	129430	14 37.4 +21 33	6.4	G5	G8 III
83273	32.1 +25 08	6.6	F8	G0 III	129537	38.0 +15 08	6.6	F2	F2 III
84497	40.6 +20 57	7.4	G5	G8 III	130603	44.0 +24 47	6.0	F5	F2 V
85586	47.7 +20 25	7.5	G5	G8 III	130948	45.8 +24 20	5.8	G0	G2 V
86419	53.3 +11 26	7.3	F2	F0 III	131473	48.7 +16 06	6.4	G0	F6 III
87178	58.2 +10 23	7.1	F2	F6 III	133161	57.9 +16 26	7.0	G0	G2 V
87776	10 02.2 +15 40	7.2	F8	G0 V	134792	15 06.7 +29 37	7.1	F5	F6 III
88009	03.7 +19 01	7.1	G5	G8 II	135101	08.3 +19 40	6.4	G5	G5 V
88355	06.2 +13 51	6.4	F5	F7 V	136138	13.9 +20 57	5.7	G5	G5 IV
96479	11 02.0 +10 44	7.3	F8	F5 IV	136176	14.1 +27 12	6.6	F8	G0 V

* *Lick Observatory Bulletin*, No. 607.

† Because there were many typographical errors in Table I as originally published, this entire paper is reprinted here. The editors regret this unfortunate occurrence, which was in no way the fault of the authors.

TABLE I (continued)

HD	α (1900) δ		m_v	Type		HD	α (1900) δ		m_v	Type	
				HD	MK					HD	MK
137510	15 ^b 21 ^m 4	+19°50'	6.3	G0	G0 V	190516	20 ^h 00 ^m 2	+15°38'	7.1	G0	G0 V
139457	33.2	+10 35	7.0	F8	F8 V	191837	06.8	+19 05	7.4	F8	F5 V
139862	35.4	+12 23	6.3	G5	G8 II	200877	21 01.1	+14 56	6.6	F5	F7 IV
140027	36.4	+16 21	6.0	G5	G8 III	201545	05.2	+18 49	7.0	F5	F8 V
140572	39.3	+23 16	7.2	F8	F8 V	201891	07.4	+17 21	7.3	F5	F7 V
140667	39.8	+11 36	7.4	F8	G0 V	205603	31.3	+14 39	6.7	G5	G8 II
159797	17 31.5	+19 54	6.7	G5	G8 III	208363	50.7	+20 23	6.9	G5	K1 III
160910	37.4	+16 00	5.6	F5	F5 V	209421	58.2	+15 18	7.2	F2	F4 V
160935	37.6	+21 34	6.7	F5	F8 IV	209601	59.5	+13 10	7.0	F5	F6 V
161149	38.8	+14 20	6.2	F5	F5 II	210211	22 03.7	+23 41	6.6	G5	G2 V
161502	17 40.8	+11 11	7.1	G5	G5 III	210264	22 04.1	+21 40	7.1	G5	G5 III
162865	48.1	+16 55	6.6	F5	F3 III-IV	210444	05.4	+22 38	7.3	F5	F6 IV
163151	49.6	+11 10	6.3	F5	F5 Vn	210460	05.5	+19 08	6.1	G0	G0 V
163840	53.2	+24 01	6.4	G0	G2 V	211152	10.1	+23 48	7.0	F5	F7 III
164043	54.2	+14 52	7.1	F8	F8 V	211460	12.1	+28 41	6.8	G5	G5 IV
164507	56.4	+15 07	6.3	G5	G5 IV	211476	12.2	+12 23	6.9	G0	G2 V
164595	56.8	+29 34	7.2	G5	G2 V	212395	18.8	+20 20	6.1	F5	F7 V
164901	58.3	+16 49	7.5	G0	F7 V	214567	34.0	+19 00	5.8	G5	G8 II
167134	18 08.8	+16 14	6.7	F5	F8 IV	215243	38.7	+10 26	6.4	F5	G8 IV
167588	10.8	+29 12	6.5	G0	F8 V	216219	46.0	+17 28	7.4	G5	G0 IIp
167944	18 12.4	+12 02	7.1	F5	F8 III	216723	22 50.1	+27 28	7.3	G5	G8 III
168874	16.9	+27 29	7.1	G5	G2 IV	217014	52.5	+20 14	5.6	G0	G5 V
170829	26.4	+20 46	6.6	G5	G8 V	217783	58.0	+22 37	7.2	F5	F5 IV
171365	29.4	+17 39	7.0	F5	F6 V	217813	58.2	+20 23	6.7	G0	G5 V
171874	32.0	+12 55	7.4	F5	F6 II	217924	59.0	+21 03	7.3	G0	G0 V
172744	36.8	+12 09	7.0	G5	K2 II	218099	23 00.2	+24 00	7.1	F5	F5 IV
174160	44.1	+23 24	6.0	F5	F8 V	218261	01.5	+19 22	6.4	F8	F7 V
176155	53.8	+17 14	5.4	F5	F8 Ib	219291	09.6	+29 13	6.4	F5	F5 IV
176973	57.6	+18 59	6.8	G5	K0 III	219477	10.9	+27 42	6.5	G5	G5 III
177082	58.1	+14 26	6.8	G0	G2 V	219779	13.4	+27 03	7.4	F5	G0 IV
						220334	23 17.8	+20 01	6.6	G0	G2 V
						221271	25.6	+16 01	7.3	G5	K1 III
						221627	28.6	+17 16	6.7	G0	G2 IV
						222249	33.9	+22 46	7.0	F5	F5 IV
						222317	34.4	+27 41	7.0	G0	G2 V
						225292	59.9	+27 07	6.6	G5	G8 II

additional stars is based on plate material obtained in 1965-66 with the same equipment, but since the exposure density was not so carefully controlled, the weight of the types assigned here is slightly lower than those of the first list. The results are given in

Table I, in which the classifications are those assigned by Harlan by the same procedures as in Paper I.

REFERENCES

Harlan, E. A. 1969, *Astron. J.* 74, 916; *Lick Obs. Bull.*, No. 605.

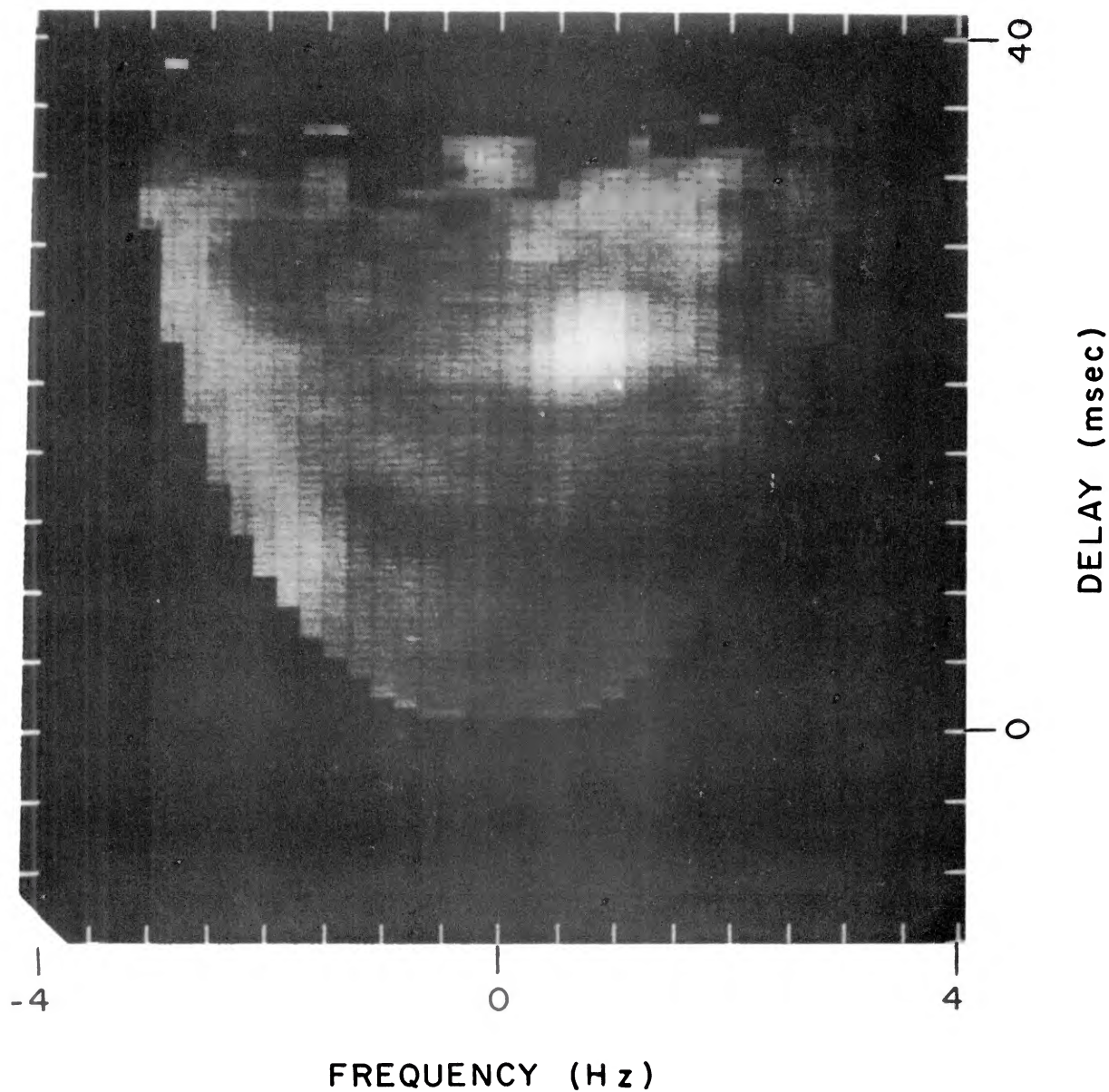


PLATE I (No. 1, Jurgens and Dyce). Delay-Doppler map of the ratio of observed power on 4 September 1967 to predicted power based on Eq. (14) using $\tau=0.05$.

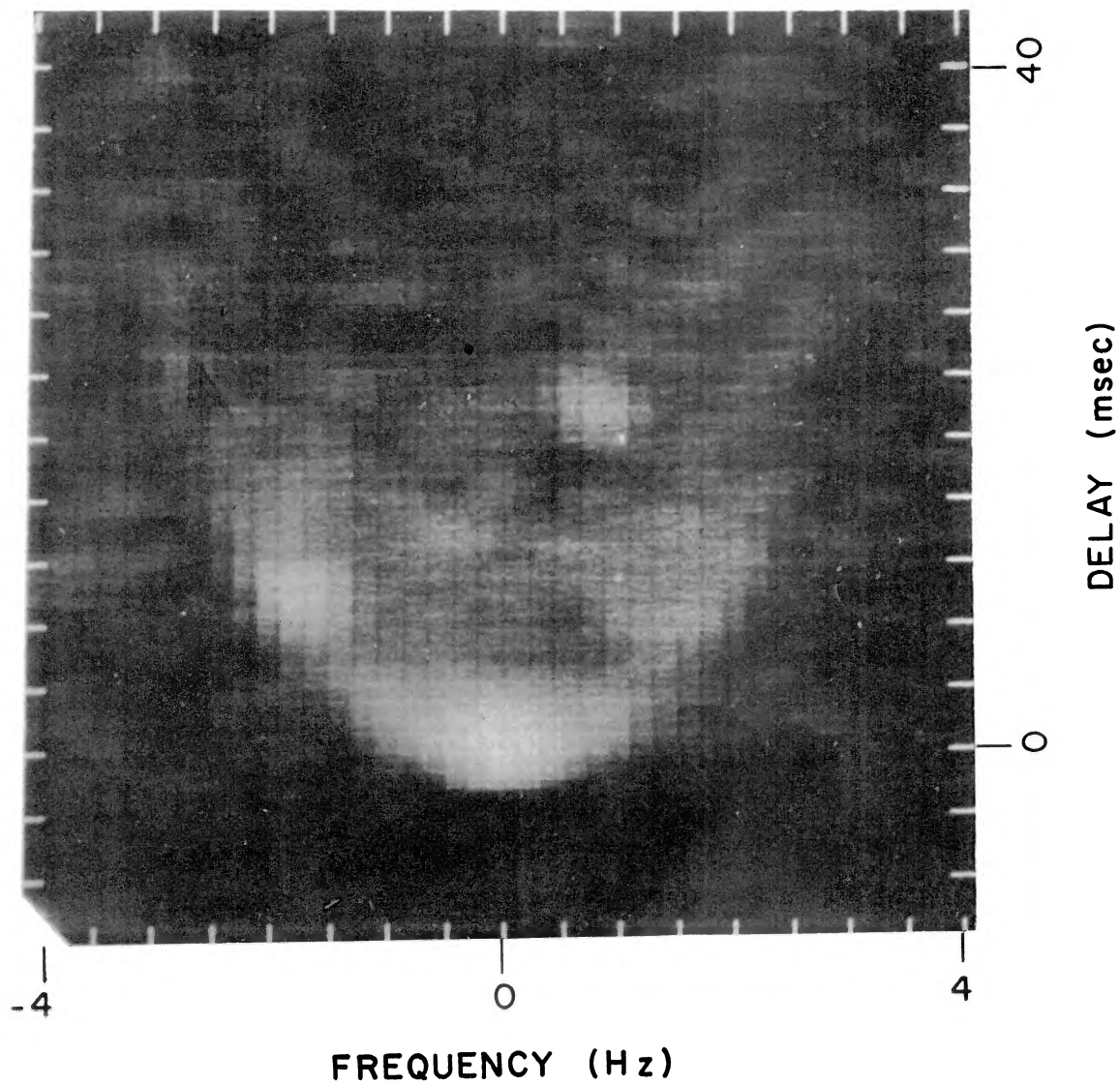


PLATE II (No. 2, Jurgens and Dyce). Delay-Doppler map of the cross-polarized power on 4 September 1967. Brightness is proportional to the $\log(P/P_{\min})$, where P_{\min} was the smallest power measurement to be displayed.

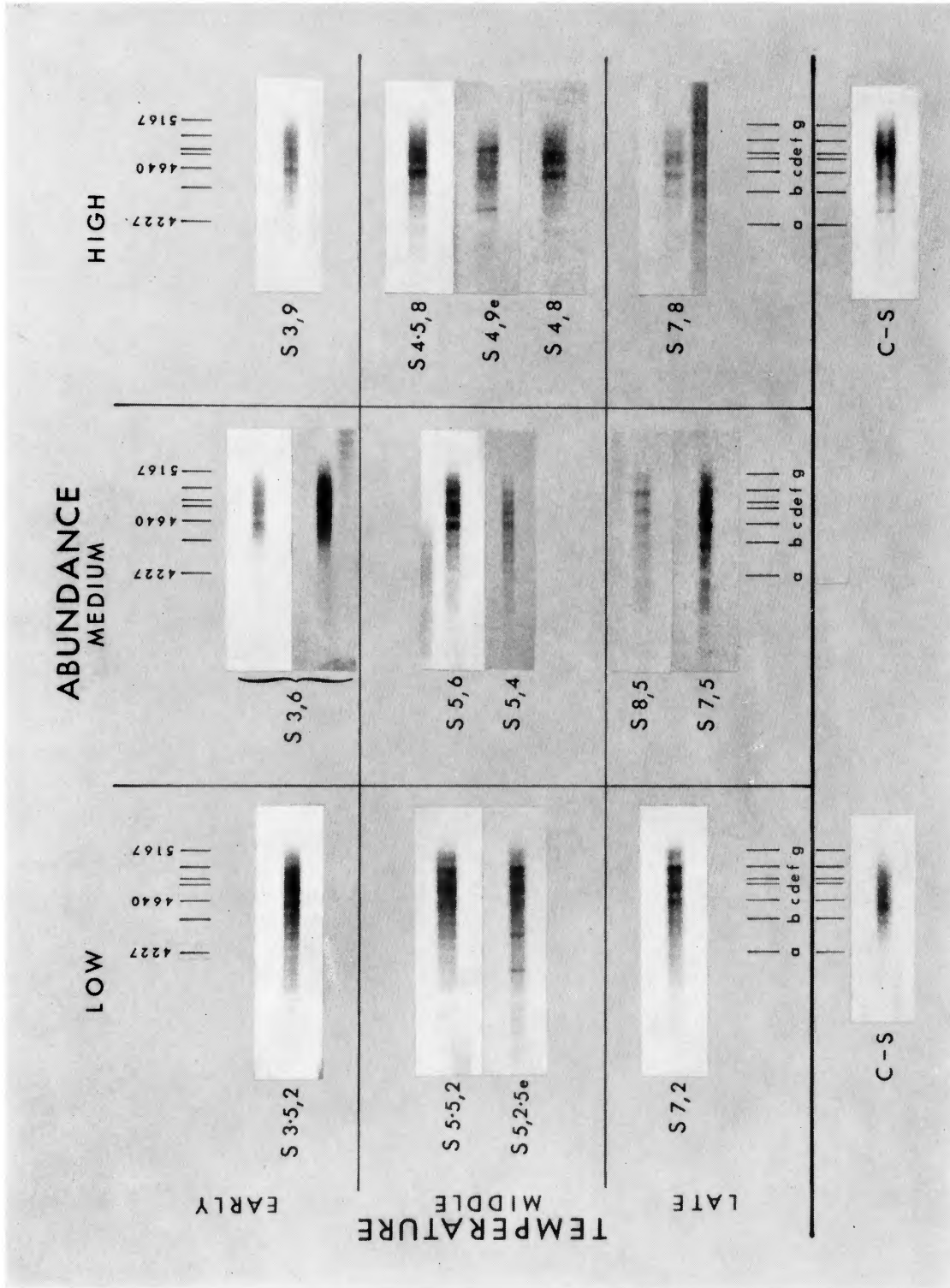


PLATE III (No. 1, Stephenson and Ross). Objective-prism spectrograms of S stars (and one weak carbon star) in the blue spectral region, original dispersion 580 Å/mm: The classifications are by Keenan (1954) except for the C-S stars, and the stars are identified in Table II. Lettered spectral features are identified in Table I.

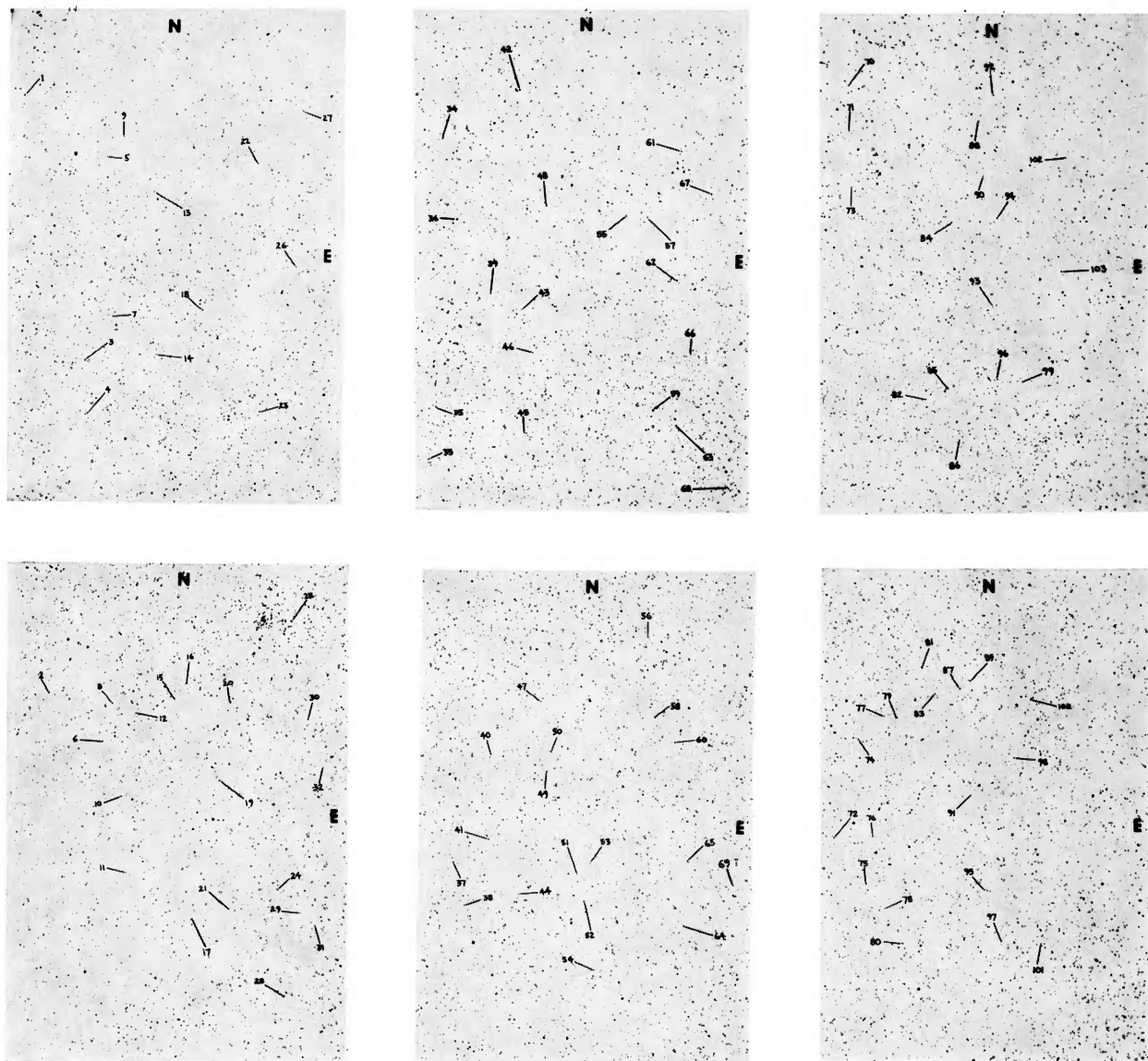



PLATE IV (No. 1, Wooden). Identification charts for the region LF 11: The scale on each of the small charts is 125 arc sec/mm. 
The center of the region is R.A. $6^{\text{h}}52^{\text{m}}5$, Dec. $+00^{\circ}51'$ (1950).

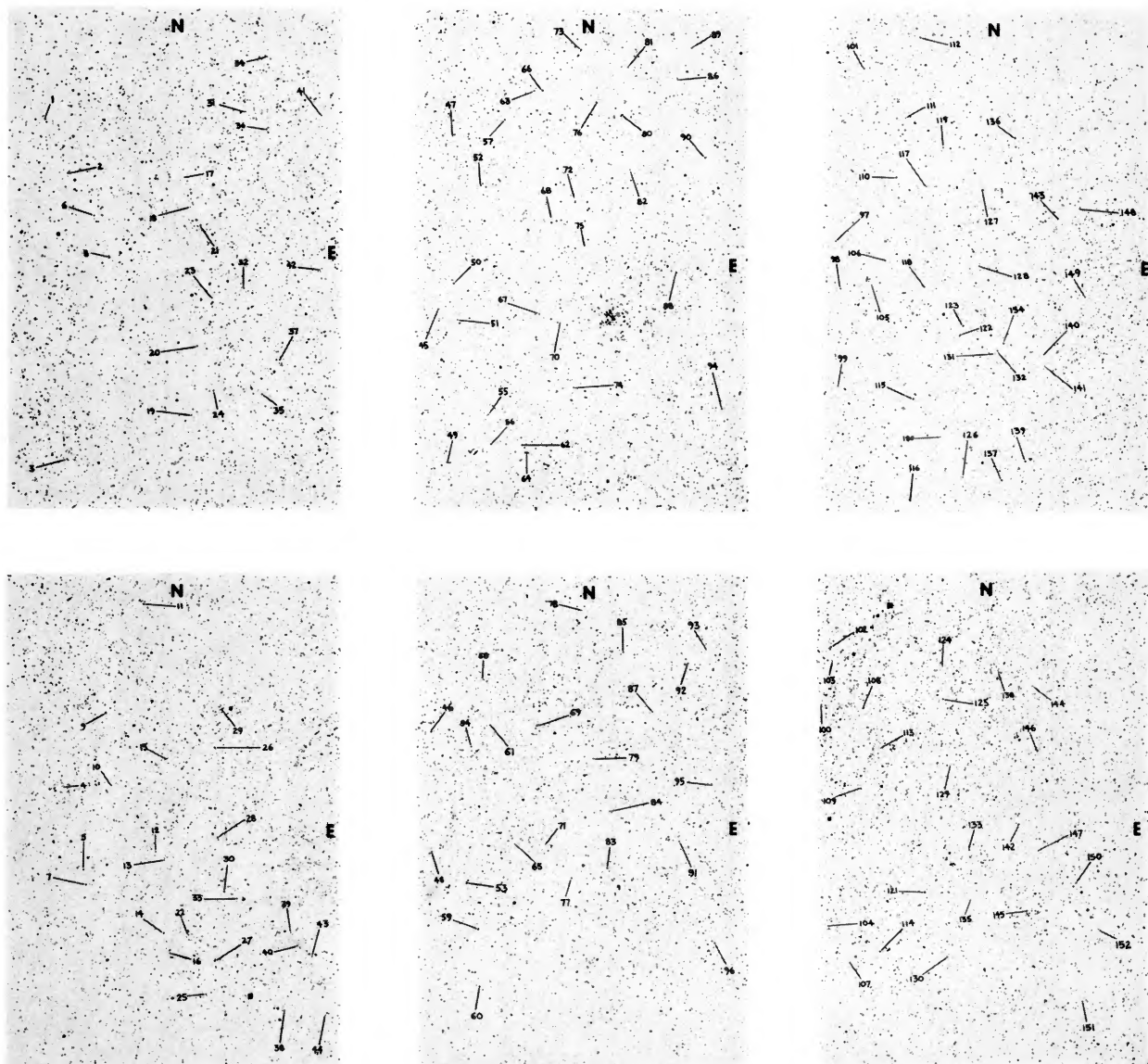


PLATE V (No. 2, Wooden). Identification charts for the region LF 12: The scale on each of the small charts is 125 arc sec/mm. The center of the region is R.A. $7^{\text{h}}41^{\text{m}}5$, Dec. $-24^{\circ}41'$ (1950).

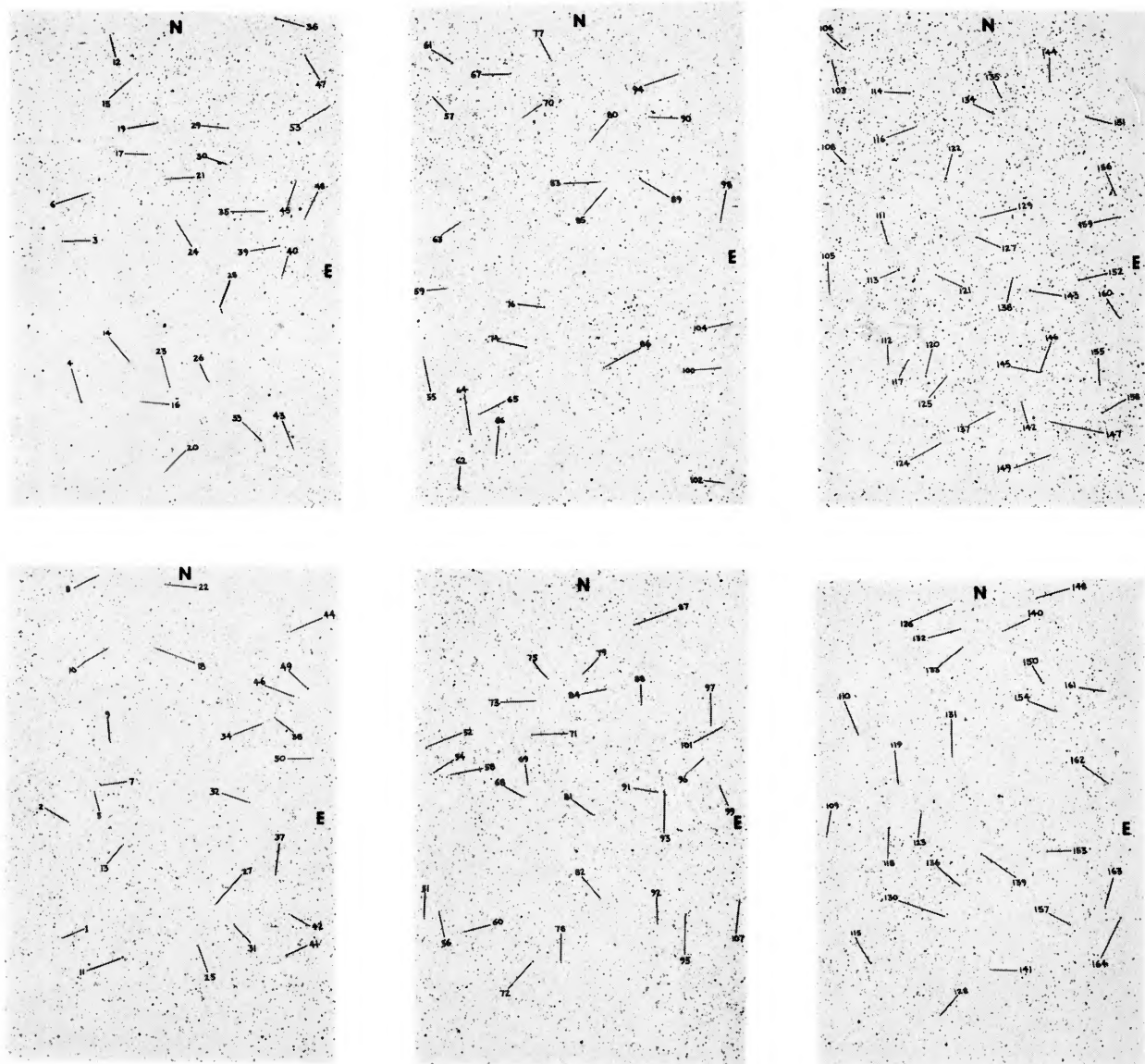


PLATE VI (No. 3, Wooden). Identification charts for the region LF 13: The scale on each of the small charts is 125 arc sec/mm. The center of the region is R.A. $10^{\text{h}}19^{\text{m}}0$, Dec. $-52^{\circ}12'$ (1950).

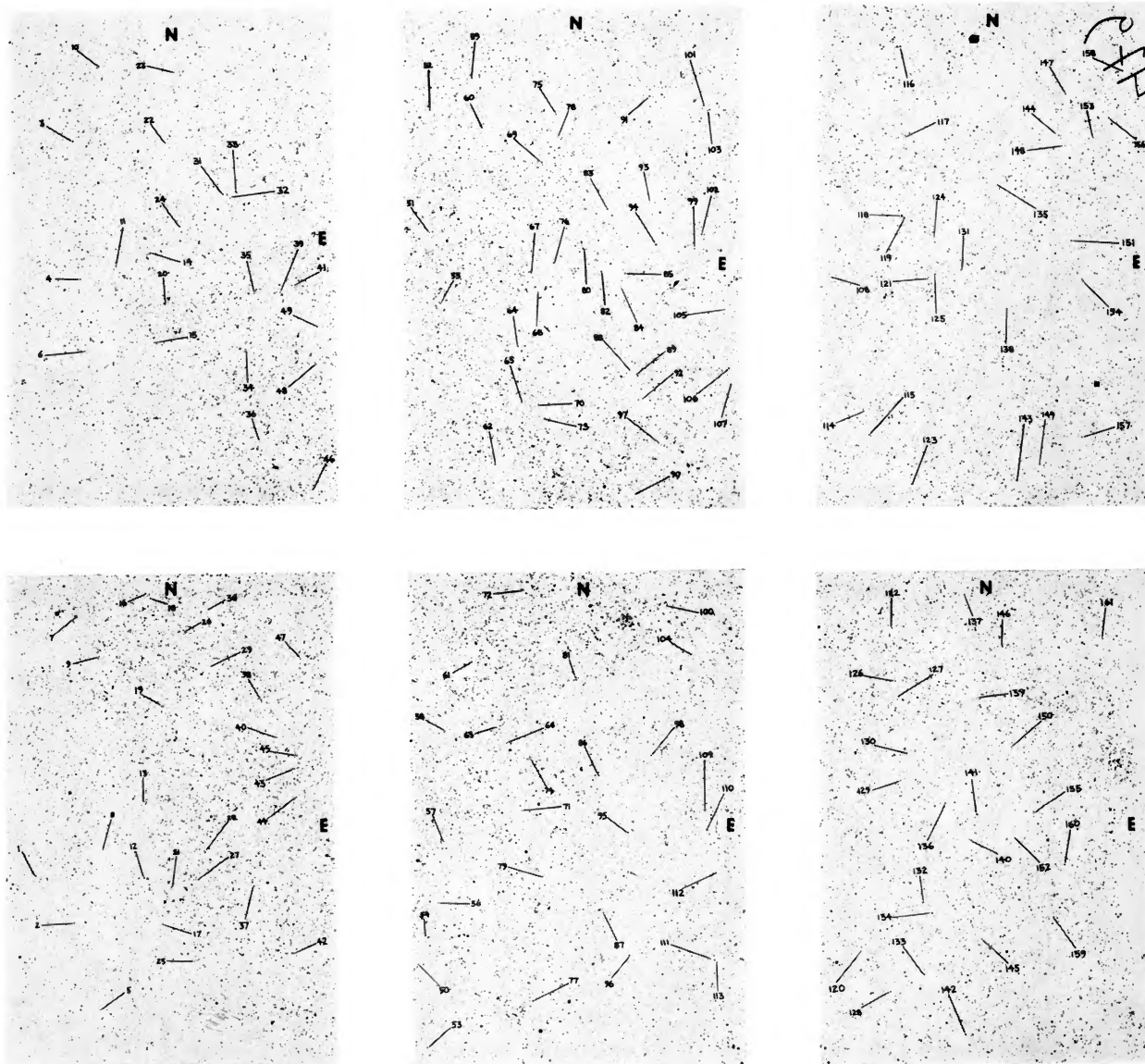


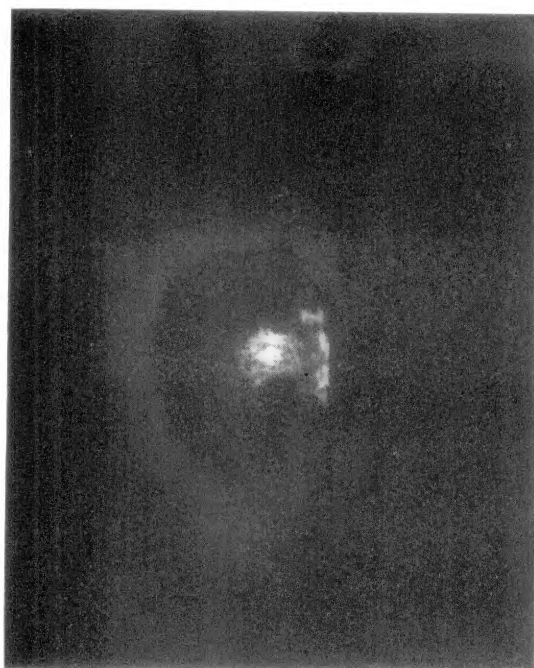
PLATE VII (No. 4, Wooden). Identification charts for the region LF 14: The scale on each of the small charts is 125 arc sec/mm. The center of the region is R.A. $12^{\text{h}}04^{\text{m}}3$, Dec. $-60^{\circ}44'$ (1950).



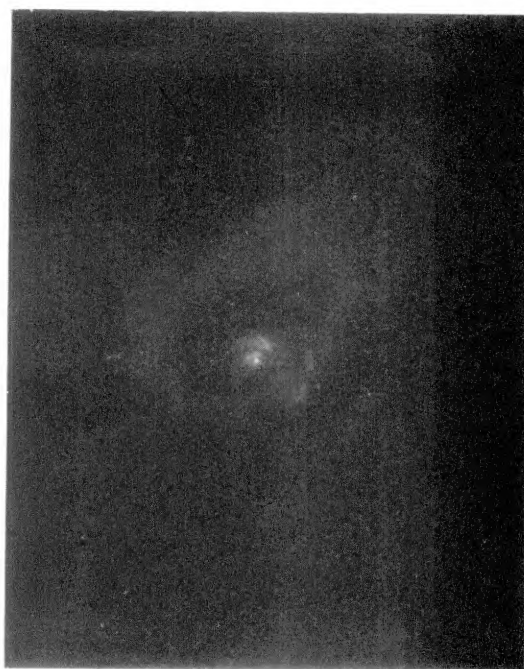
-14.6 km/sec
(a)



-3.2 km/sec
(b)



+8.2 km/sec
(c)



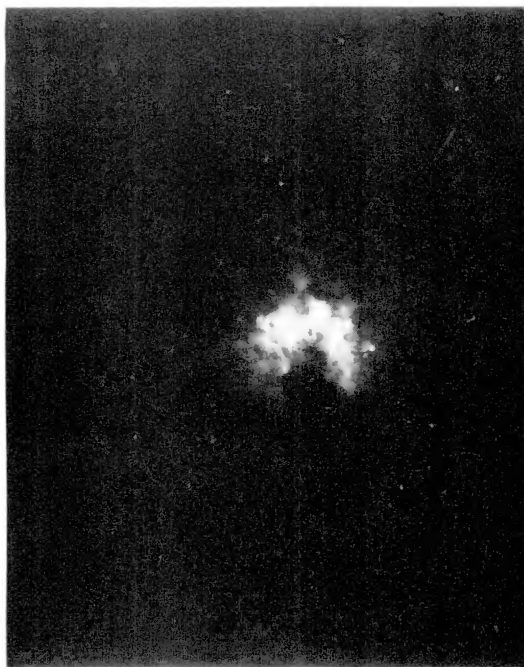
+19.6 km/sec
(d)

PLATE VIII (No. 1, Fisher and Williamson). Narrow-band filtergrams of the Orion Nebula in the wavelength region 6563-6583 Å.



+31.0 km/sec

(e)



+42.4 km/sec

(f)



+53.0 km/sec

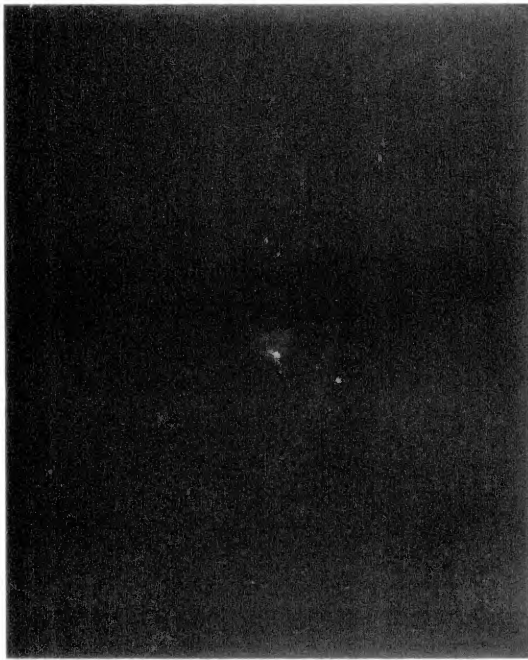
(g)



+65.2 km/sec

(h)

PLATE VIII (No. 1, Fisher and Williamson) *(continued)*.



+88.0 km/sec

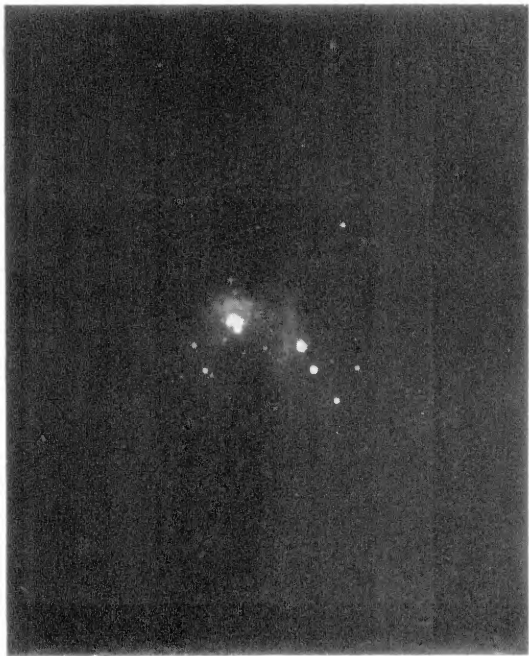
(i)



(j)



(k)



(l)

PLATE VIII (No. 1, Fisher and Williamson) (*continued*).

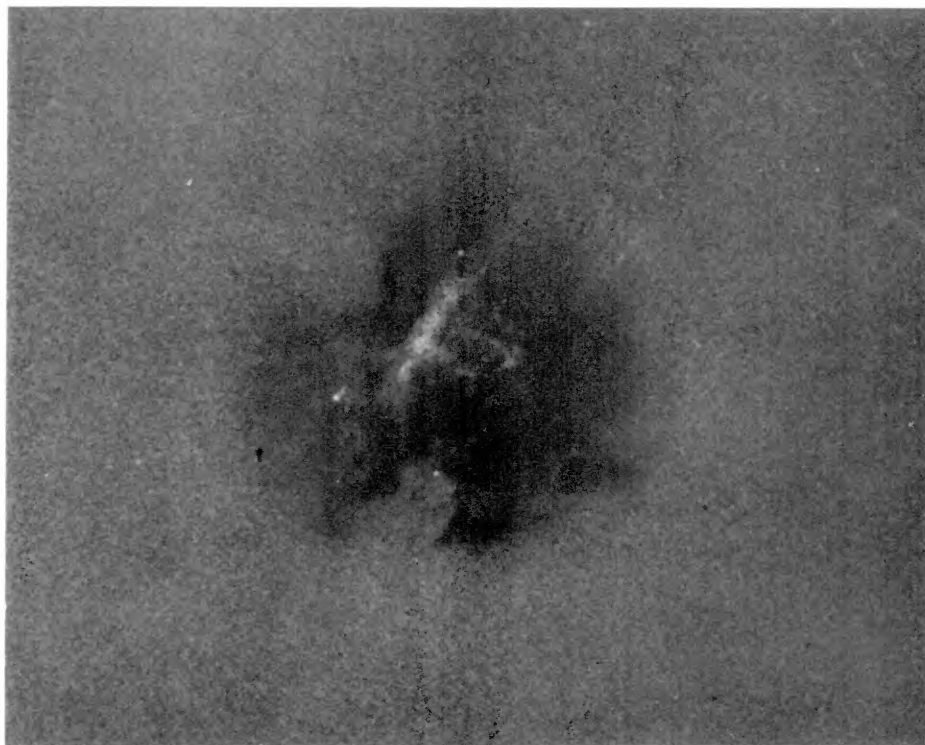


PLATE IX. (No. 2, Fisher and Williamson). Plate VIII (b) and (f) E and W are exchanged relative to Plate VIII (a)-(l).

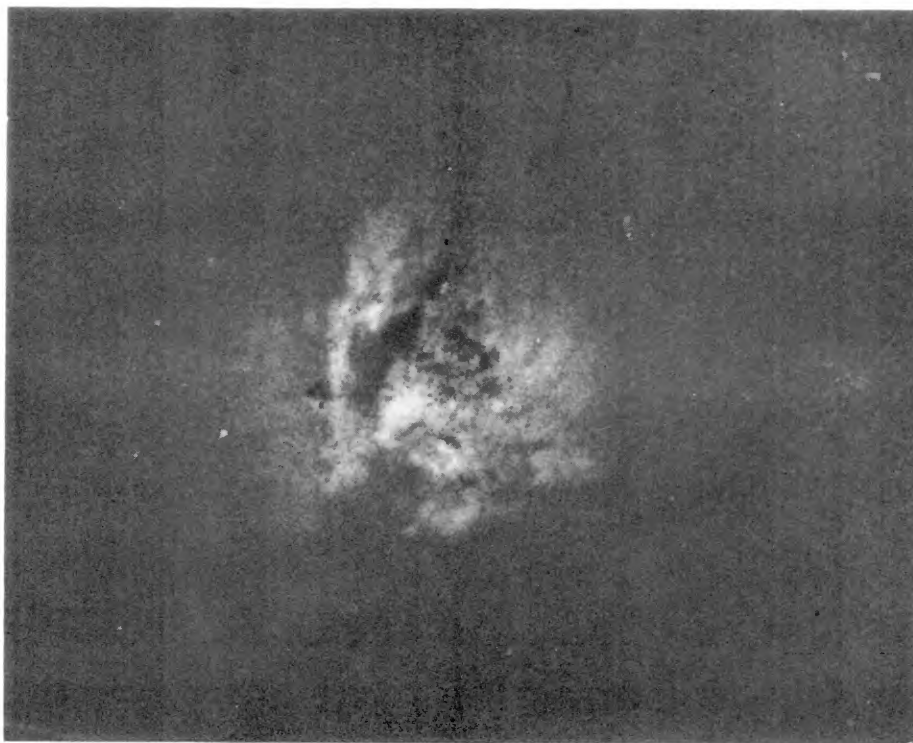


PLATE X. (No. 3, Fisher and Williamson). Plate VIII(f) and (b) E and W are exchanged relative to Plate VIII (a)-(l).